# ERTH 491-01 / GEOP 572-02 Geodetic Methods

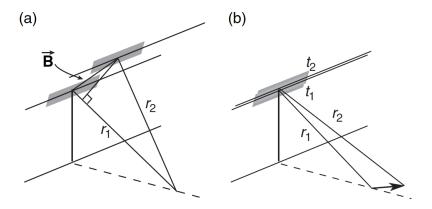
### - Lecture 12: InSAR - Making the Interferogram –

Ronni Grapenthin rg@nmt.edu MSEC 356 x5924

September 30, 2015

#### **Difference between InSARs**

Topo: look at same thing from 2 views (SRTM) Deformation: look at same thing from same point and see whether it moved



Simons and Rosen, 2007

Make interferogram from 2 Single Look Complex images (images are in radar coordinates: range  $\rho$ , azimuth *a*):

- 1 align reference and repeat images to sub-pixel accuracy
- 2 multiply complex images (SLC) to form complex interferogram
- 3 extract phase:  $\phi_2 \phi_1 = \arctan \frac{lm}{Re}$

- take 100s of small sub-patches (e.g.  $64\times 64)$  from master and slave
- 2D cross correlation of patch pairs
- determine 6-parameter affine transformation to align slave to master image
- affine: parallel remains, straight remains, points preserved

#### Making an Interferogram: Step 2 - Multiply

Complex number of each pixel, C(x), in terms of amplitude, A(x), and phase,  $\phi(x)$ :

$$C(x) = A(x)e^{i\phi(x)}$$

with  $x = (\rho, a)$ 

#### Making an Interferogram: Step 2 - Multiply

Complex number of each pixel, C(x), in terms of amplitude, A(x), and phase,  $\phi(x)$ :

$$C(x) = A(x)e^{i\phi(x)}$$

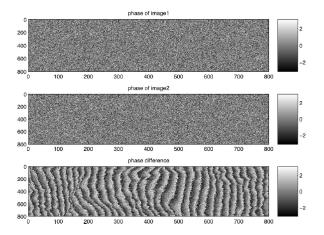
with  $x = (\rho, a)$ 

Multiply (pixel by pixel, note complex conjugate!):

$$C_2 C_1^* = A_2 A_1 e^{i(\phi_2 - \phi_1)} \\ = Re(x) + i Im(x)$$

#### Making an Interferogram: Step 3 - Get Phase

 $\phi_2 - \phi_1 = \arctan \frac{lm}{Be}$ 



What's in the phase?

$$\phi = \mathbf{E} + \phi_{topo} + \mathbf{D} + \epsilon_{orbit} + \mathbf{I} + \mathbf{T} + \epsilon$$

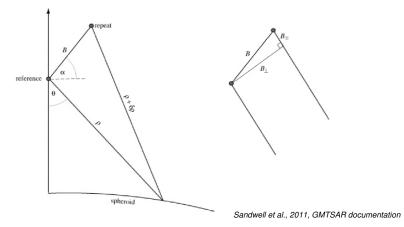
where:

- E: earth curvature (almost planar, known)
- $\phi_{topo}$ : topographic phase (broad spectrum)
- D: surface deformation (unknown, we want to know!)
- $\epsilon_{orbit}$ : orbit error (almost a plane, mostly known)
- I: Ionospheric Delay (plane or 4 km wavelength waves!)
- T: Tropospheric Delay (power law, unknown)
- $\epsilon$ : phase noise (white, unknown)

#### EarthShape = curvature + topography

#### Phase due to Earth Curvature

#### Repeat-pass interferometry geometry:



assume parallel paths:

$$egin{array}{rcl} B_{\parallel} &=& B \sin( heta-lpha) \ B_{\perp} &=& B \cos( heta-lpha) \end{array}$$

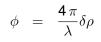
Issue?

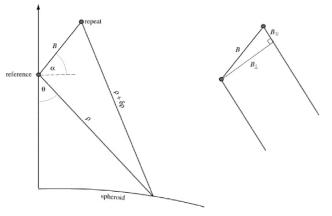
#### Issue?

- $\theta$  (look angle) changes across path
- single  $B_{\perp}$  for combination is approximate!

### Phase due to Earth Curvature

Phase difference  $\phi$ relates to range difference  $\delta \rho$ :

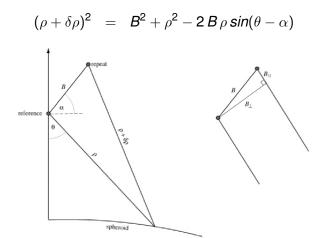




Sandwell et al., 2011, GMTSAR documentation

#### Phase due to Earth Curvature

Law of Cosines:



Sandwell et al., 2011, GMTSAR documentation

Law of Cosines:

$$(\rho + \delta \rho)^2 = B^2 + \rho^2 - 2 B \rho \sin(\theta - \alpha)$$

Algebra:

$$\begin{aligned} (\rho + \delta\rho)^2 &= B^2 + \rho^2 - 2 B \rho \sin(\theta - \alpha) \\ 2\rho\delta\rho + \delta\rho^2 &= B^2 - 2 B \rho \sin(\theta - \alpha) \\ 2\rho(\delta\rho + \frac{\delta\rho^2}{2\rho}) &= 2\rho(\frac{B^2}{2\rho} - B\sin(\theta - \alpha)) \\ \delta\rho &= -B\sin(\theta - \alpha) + \frac{1}{2\rho}(B^2 + \delta\rho^2) \end{aligned}$$

#### Phase due to Earth Curvature

Law of Cosines (after algebra):

$$\delta \rho = -B\sin(\theta - \alpha) + \frac{1}{2\rho}(B^2 + \delta \rho^2)$$

Assume  $\delta \rho << \rho$ :

$$\delta 
ho = rac{B}{2
ho} - B \sin( heta - lpha)$$

Assume  $B << \rho$ :

$$\delta \rho = -B \sin(\theta - \alpha)$$

Plug into phase difference eqn:

$$\phi = \frac{-4 pi}{\lambda} B \sin(\theta - \alpha)$$

$$\phi = \frac{-4 \, pi}{\lambda} B \sin(\theta - \alpha)$$

- Making some more assumptions we could get phase correction based on  $\delta\rho<<<\rho$  assumption
- instead, use phase expression to approximate higher order terms
- need ot think about phase change wrt to range first

Differentiate w.r.t. range:

$$\begin{array}{lll} \frac{\delta\phi}{\delta\rho} & = & \frac{-4\,\textit{pi}}{\lambda}\textit{B}\cos(\theta-\alpha)\frac{\delta\theta}{\delta\rho} \\ & = & \frac{-4\,\textit{pi}}{\lambda}\textit{B}_{\perp}\frac{\delta\theta}{\delta\rho} \end{array} \end{array}$$

Depends on  $B_{\perp}$  and derivative of look angle wrt range:  $\frac{\delta\theta}{\delta a}$ 

# Look angle wrt range

Need  $\frac{\delta\theta}{\delta\rho}$ :

- assume locally spherical Earth
- adjust local radius,  $r_e(\psi)$ , of the Earth using WGS-84 ellipsoid
- large differences between spherical and elliptical Earth for long SAR swaths!
- would give fringes (artifacts) that depend on baseline
- get correction term later!

$$r_{e}(\psi) = \left(\frac{\cos^{2}(\psi)}{a^{2}} + \frac{\sin^{2}(\psi)}{c^{2}}\right)^{-\frac{1}{2}}$$

•  $\psi$ : latitude

- a: equatorial radius 6378 km
- c: polar radius 6357 km

$$\delta 
ho = -B \sin(\theta - \alpha) + \frac{1}{2\rho} (B^2 + \delta \rho^2)$$

... keep higher order term, but use  $\delta \rho = -B \sin(\theta - \alpha)$  found above

$$\begin{split} \delta\rho &= -B\sin(\theta-\alpha) + \frac{B^2}{2\rho}(1-\sin^2(\theta-\alpha)) \\ &= -B\sin(\theta-\alpha) + \frac{B^2}{2\rho}(\cos^2(\theta-\alpha)) \end{split}$$

... could iterate to get even more accurate ...

ratio of first and second order terms:

$$\begin{array}{ll} \textit{ratio} & = & \displaystyle \frac{B\cos^2(\theta-\alpha)}{2\rho\sin(\theta-\alpha)} \\ & = & \displaystyle \frac{3\,B}{4\,\rho} \end{array} \end{array}$$

... if we assume  $\alpha = 0, \ \theta = 30^{\circ}$ 

 map topography from lat, lon, height to radar coordinates and topography over range, azimuth t(ρ, a)

- map topography from lat, lon, height to radar coordinates and topography over range, azimuth t(ρ, a)
- · read row of data from reference and repeat image

- map topography from lat, lon, height to radar coordinates and topography over range, azimuth t(ρ, a)
- read row of data from reference and repeat image
- use precise spacecraft orbit to for reference image to get b, B,  $\alpha$

- map topography from lat, lon, height to radar coordinates and topography over range, azimuth t(ρ, a)
- read row of data from reference and repeat image
- use precise spacecraft orbit to for reference image to get *b*, *B*,  $\alpha$
- interpolate topography to each range pixel get look angle from:

$$heta_{
ho,a} = cos^{-1} \left[ rac{(b^2 + rho^2 - (r_e + t(
ho, a)))^2}{2
ho b} 
ight]$$

- map topography from lat, lon, height to radar coordinates and topography over range, azimuth t(ρ, a)
- read row of data from reference and repeat image
- use precise spacecraft orbit to for reference image to get *b*, *B*,  $\alpha$
- interpolate topography to each range pixel get look angle from:

$$\theta_{
ho,a} = \cos^{-1}\left[rac{(b^2 + rho^2 - (r_e + t(
ho, a)))^2}{2
ho b}
ight]$$

• with look angle for each range pixel, calculate phase correction for repeat image:

$$\phi_{\rho,a} = -\frac{4\pi B}{\lambda} sin(\theta_{\rho,a}) - \alpha) + \frac{2\pi B^2}{\lambda \rho} cos^2(\theta_{\rho,a} - \alpha)$$

- map topography from lat, lon, height to radar coordinates and topography over range, azimuth t(ρ, a)
- read row of data from reference and repeat image
- use precise spacecraft orbit to for reference image to get *b*, *B*,  $\alpha$
- interpolate topography to each range pixel get look angle from:

$$\theta_{
ho,a} = \cos^{-1}\left[rac{(b^2 + rho^2 - (r_e + t(
ho, a)))^2}{2
ho b}
ight]$$

• with look angle for each range pixel, calculate phase correction for repeat image:

$$\phi_{\rho,a} = -\frac{4\pi B}{\lambda} sin(\theta_{\rho,a}) - \alpha) + \frac{2\pi B^2}{\lambda \rho} cos^2(\theta_{\rho,a} - \alpha)$$

multiply C<sub>2</sub> C<sup>\*</sup><sub>1</sub>

- map topography from lat, lon, height to radar coordinates and topography over range, azimuth  $t(\rho, a)$
- read row of data from reference and repeat image
- use precise spacecraft orbit to for reference image to get *b*, *B*,  $\alpha$
- interpolate topography to each range pixel get look angle from:

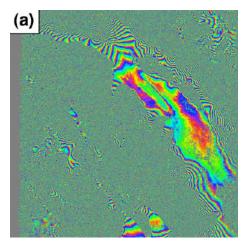
$$heta_{
ho,a} = cos^{-1} \left[ rac{(b^2 + rho^2 - (r_e + t(
ho, a)))^2}{2
ho b} 
ight]$$

• with look angle for each range pixel, calculate phase correction for repeat image:

$$\phi_{
ho,a} = -rac{4\pi B}{\lambda} sin( heta_{
ho,a}) - lpha) + rac{2\pi B^2}{\lambda 
ho} cos^2( heta_{
ho,a} - lpha)$$

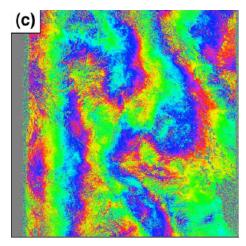
- multiply C<sub>2</sub> C<sup>\*</sup><sub>1</sub>
- extract phase difference  $\phi_2 \phi_1 = arctan(\frac{lm}{Re})$

1.95 km baseline Interferogram, no topo removed (120 fringes need removal):



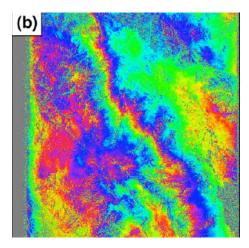
1.95 km baseline Interferogram,

topography correction using approximate formulas:



1.95 km baseline Interferogram,

topography correction using exact(er) formulas:



1.95 km baseline Interferogram,

Difference between exact and approx. formulas = 0.6 m ramp

