

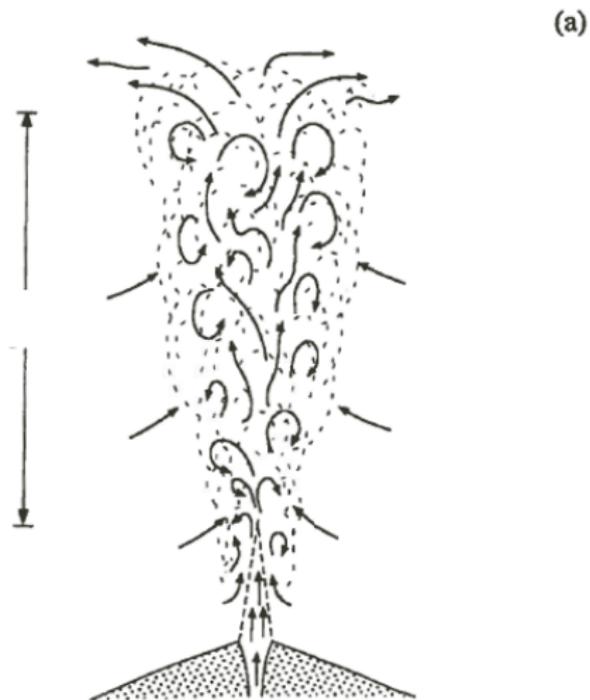
**ERTH 456 / GEOL 556**  
**Volcanology**

**– Lecture 13: Volcanic Plumes II –**

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hours: TR 3-4PM or appt.

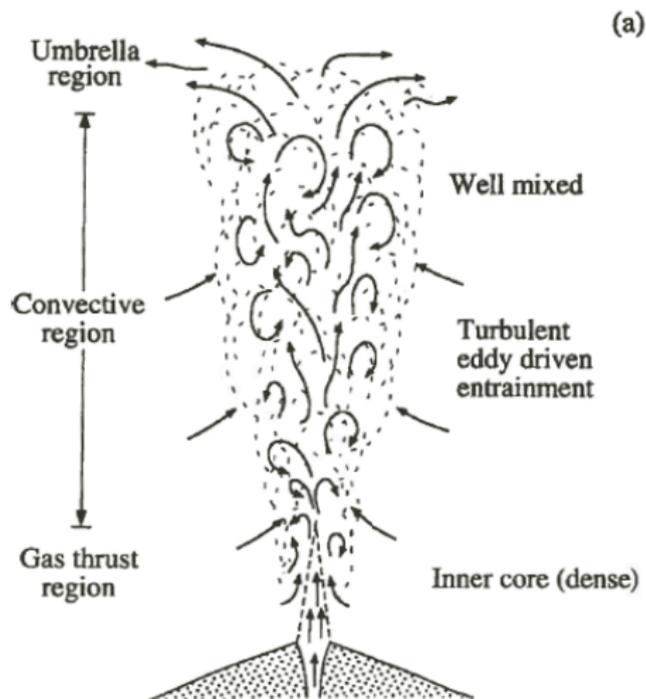
October 07, 2016

# Plumes - Densities



*Sparks et al., 1997*

# Plumes - Densities



Sparks et al., 1997

# Density Variations in Eruptive Mix

Density of mixture ( $\beta$ ) given by:

$$\frac{1}{\beta} = \frac{1-n}{\sigma} + \frac{n}{\rho}$$

$n, \rho$ : mass fraction & density of gas

$\sigma$ : density of pyroclasts

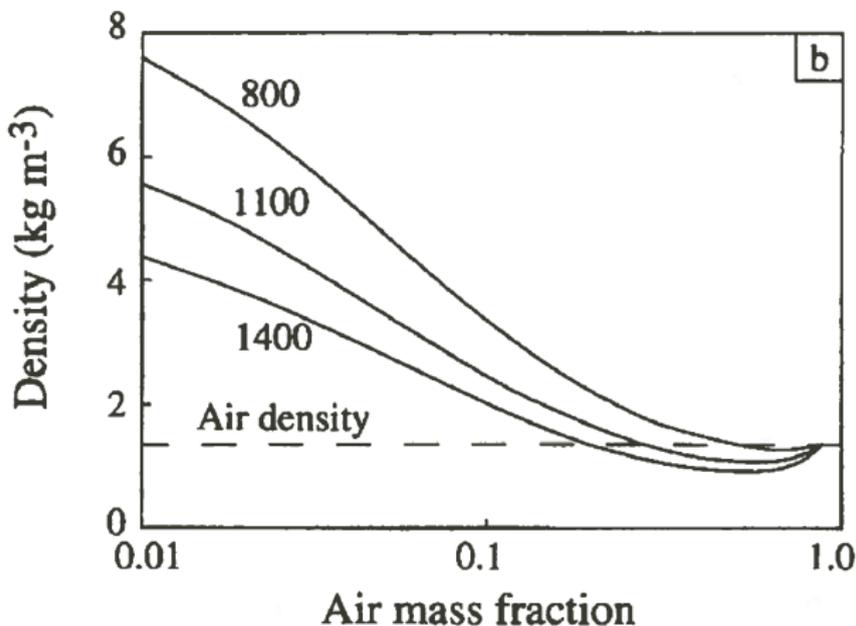
assume gas phase behaves as perfect gas:

$$\rho = \frac{P}{RT}$$

$P, T$ : pressure & Temperature of mixture

$R$ : gas constant, average if gaseous components: air= $285 \text{ Jkg}^{-1} \text{ K}^{-1}$ ,  
 $\text{CO}_2 = 185 \text{ Jkg}^{-1} \text{ K}^{-1}$ , water vapor= $460 \text{ Jkg}^{-1} \text{ K}^{-1}$

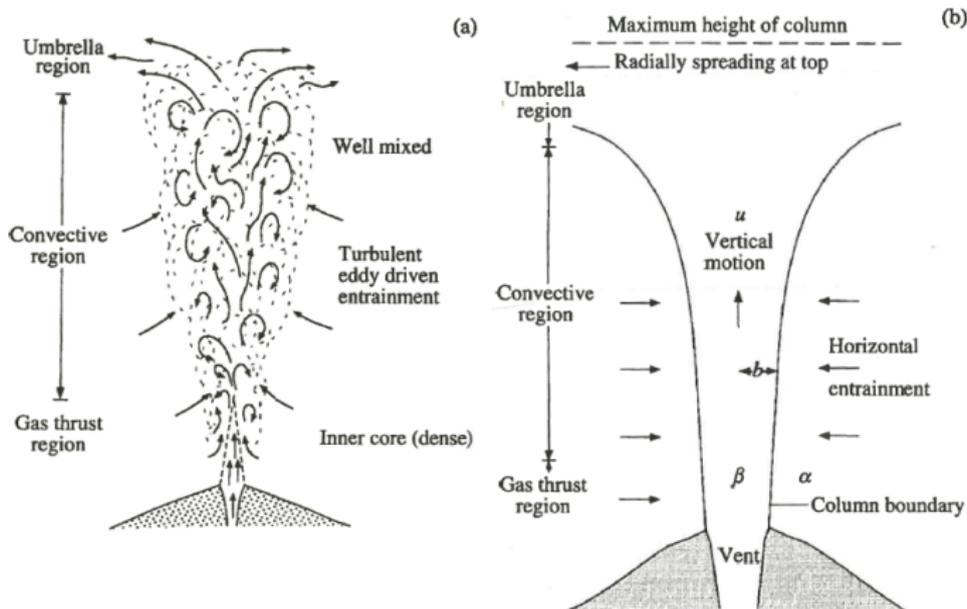
# Density Variations in Eruptive Mix



*Sparks et al., 1997*

Density of mixture (entrained air, pyroclasts, volatiles) function of entrained air; three eruption temperatures given in Kelvin & constant water 3%

# Density Variations in Eruptive Mix

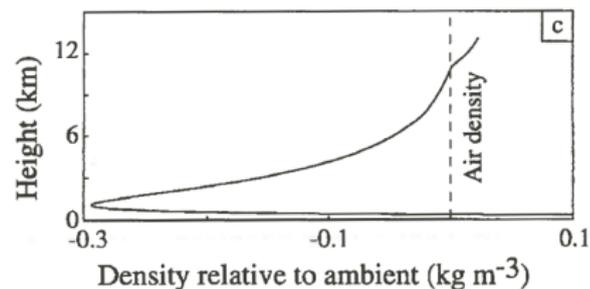
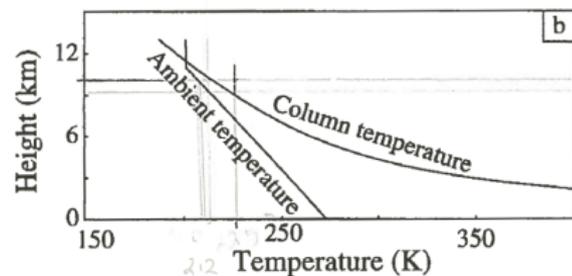
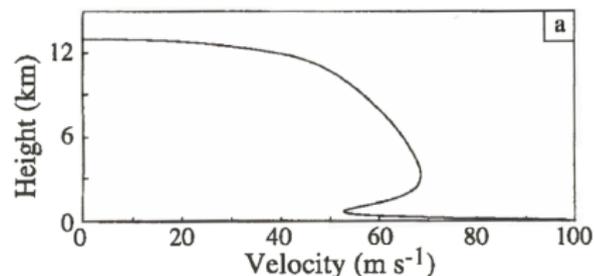


Sparks et al., 1997

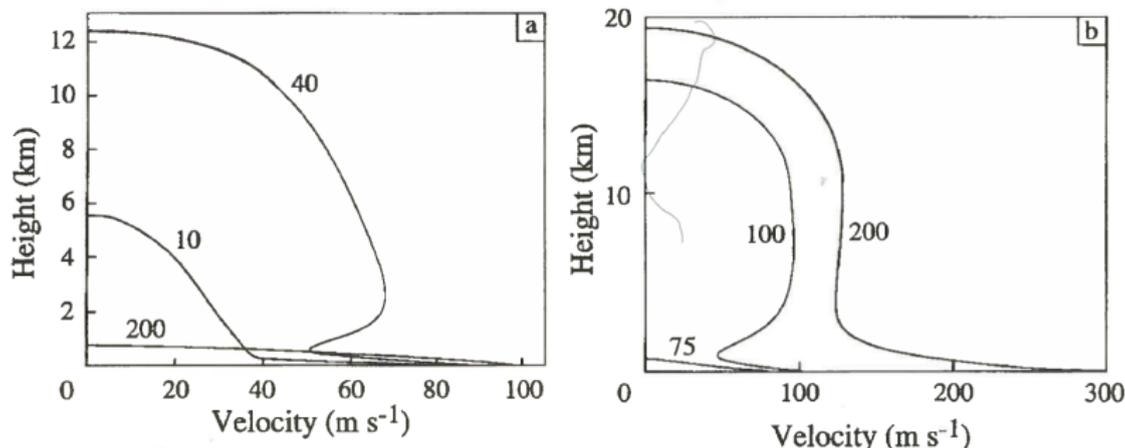
entrainment coefficients: jet  $\approx 0.06$ ; buoyant plume  $\approx 0.09$  (more efficient; other models exist)

# Density & Temperature Variations

- initial radius: 50 m
- initial velocity: 100 m /s
- eruption temperature: 1000 K
- initial mixtures 3% water (mass fraction)

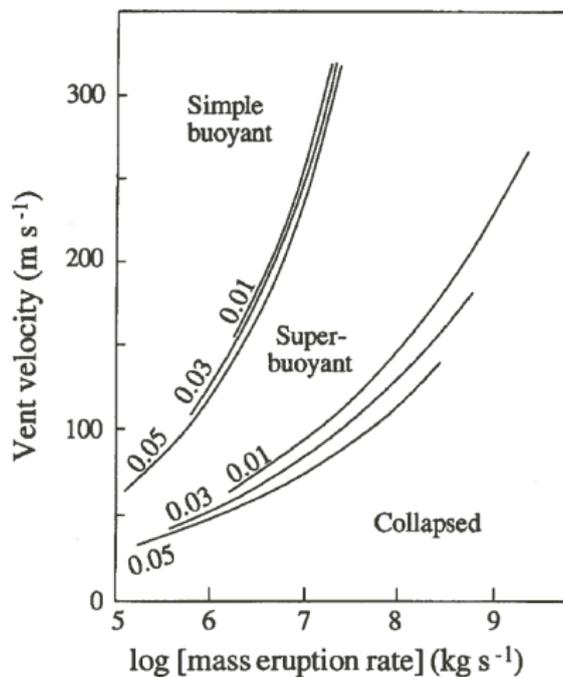


# Velocities vs. Vent Radii



**Figure 4.5** Variation of the velocity in the column as a function of the height. Curves are shown for (a) three initial radii, 10, 40 and 200 m with eruption velocity of  $100 \text{ m s}^{-1}$ , and (b) three eruption velocities 200, 100 and  $75 \text{ m s}^{-1}$  with a radius of 100 m. The mass fraction of water is 0.03 and the eruption temperature 1000 K. With the larger initial radius (a) or smaller eruption velocity (b) the material takes longer to entrain sufficient fluid to become buoyant, eventually leading to collapse in the case of the 200 m initial radius (a) and  $75 \text{ m s}^{-1}$  initial velocity (b). The 10 m vent radius (a) and  $200 \text{ m s}^{-1}$  eruption velocity (b) lead to a monotonically decaying velocity profile, since the material becomes buoyant rapidly. However, the 40 m vent radius (a) leads to a non-monotonic velocity profile, because the column entrains ambient air more slowly, and so the velocity falls off dramatically before the material becomes buoyant. A column with this non-linear velocity profile is referred to as superbuoyant. After Bursik and Woods (1991)

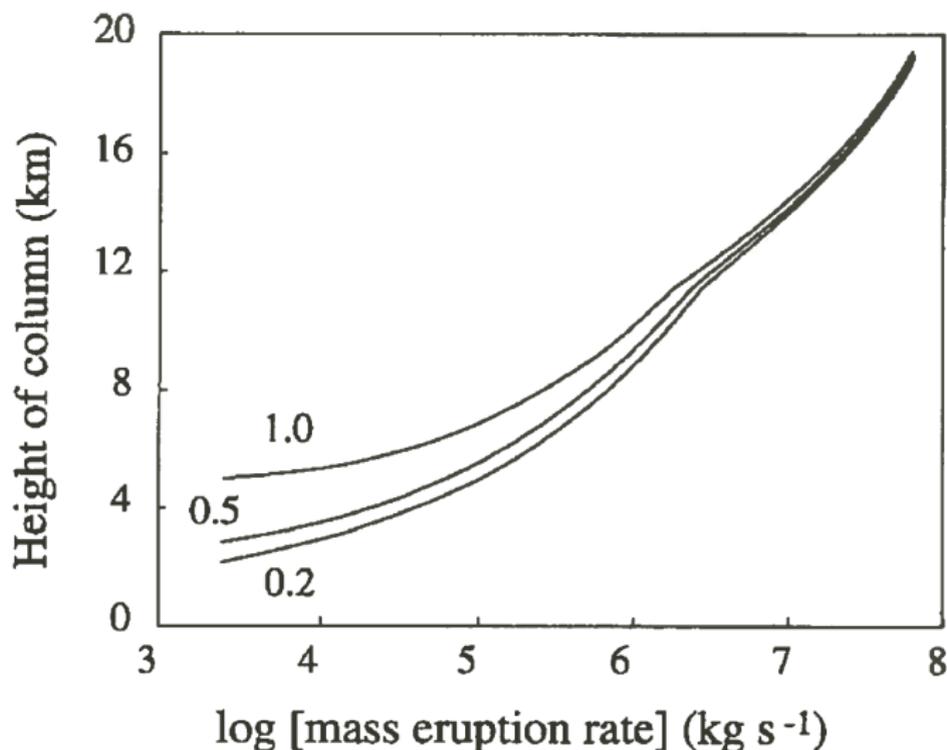
# Eruption Regimes: Velocities vs. Vent Radii



Sparks et al., 1997

Solid curves are labeled with initial mass fraction of water

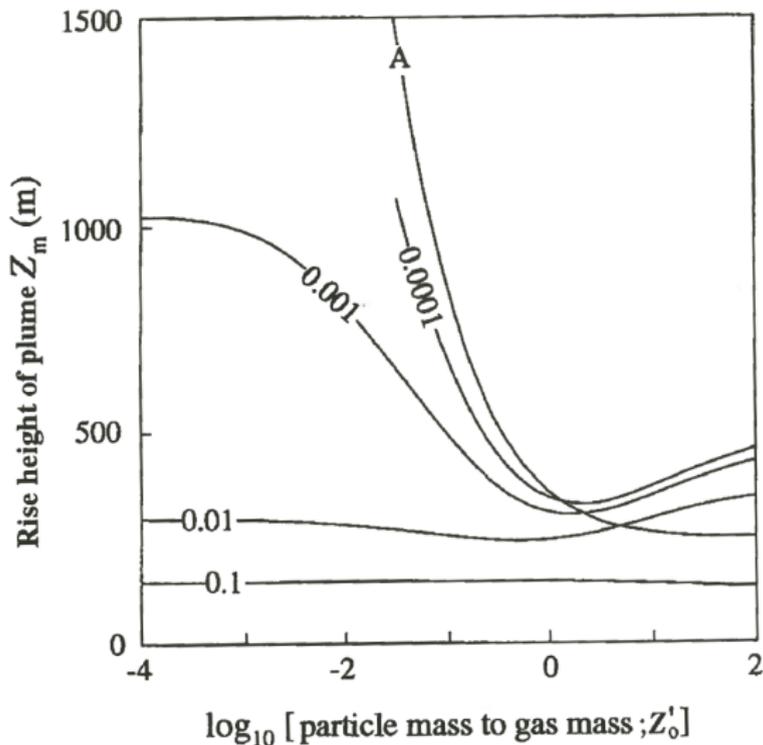
# Humidity - Vapor entrainment



*Sparks et al., 1997*

Solid curves are labeled with different relative humidities

# Jet Rise Heights - Vapor entrainment



Sparks et al., 1997

10 m vent diameter with 100 m/s initial velocity, curves for different particle radii in meters

What's the major difference? R. Grapenthin

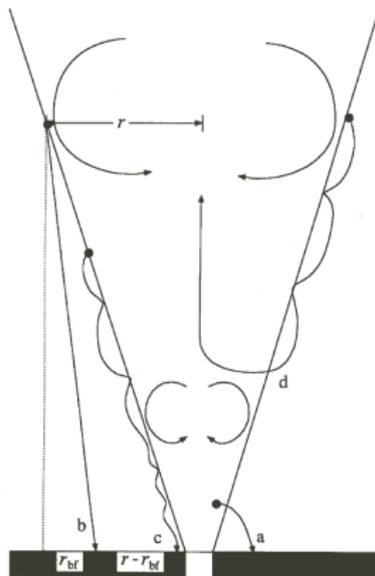


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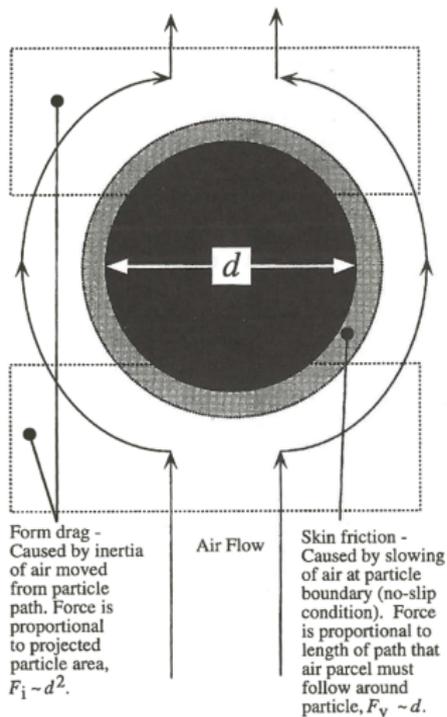
board work.

# Sedimentation Model



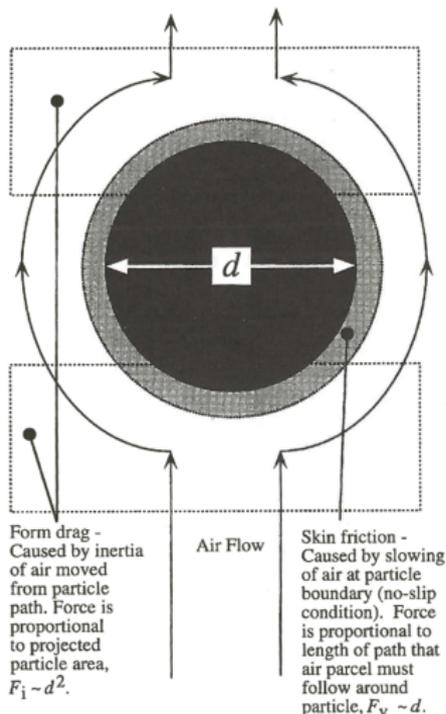
**Figure 14.8** Diagram to illustrate sedimentation processes from the margins of a plume or jet. Particles can follow a variety of paths once they fall from a plume, here shown schematically (from Ernst *et al.* 1996a). Particle "a" follows a ballistic path to the ground. Particle "b" hits the ground at a distance,  $r - r_{bf}$ , that is closer to the plume centreline than is the distance at which it fell from the plume, because of the inward flow of ambient fluid into the plume. The effect of the backflow on particle "c" is sufficiently strong that the particle is almost re-entrained, although the particle is too large to remain in the plume. Particle "d" is simply re-entrained

# Particle Settling

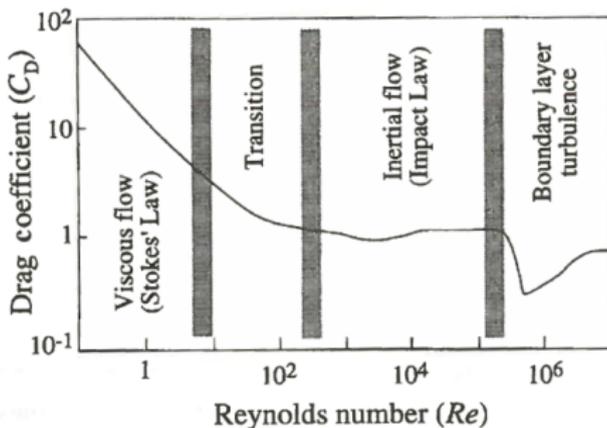


*Sparks et al., 1997*

# Particle Settling



Sparks et al., 1997



Sparks et al., 1997

- Reynold's number indicates turbidity of flow
- ratio of inertial force ( $F_i$ , enhances turbulence) to viscous force ( $F_v$ , suppresses turbulence)
- $Re = F_i/F_v = \frac{V_t d}{\nu}$
- $d$  particle diameter,  $V_t$  terminal velocity,  $\nu$  fluid kinematic viscosity (dynamic viscosity / density)

# Particle Settling - Terminal Velocity

- Terminal velocity:  $V_t = \left( \frac{4}{3} \frac{d(\sigma - \rho)g}{C_D \rho} \right)$
- $\sigma$  particle density,  $\rho$  ambient fluid density (negligible),  $g$  acc gravity,  $C_D$  drag coeff

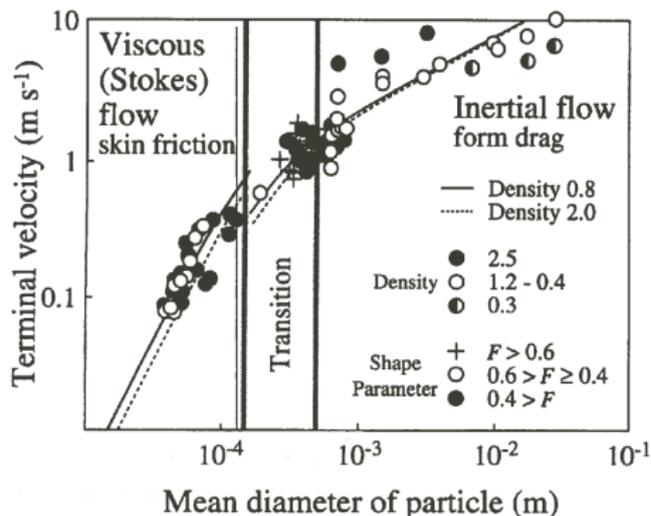
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- particles are irregular, introduce shape factor  $F$  to determine drag:
- $F = (b_p + c_p)/2a_p$  with  $a_p > b_p > c_p$  principal axes of particle

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- $C_D = \frac{24}{Re} F^{-0.32} + 2\sqrt{1.07 - F}$

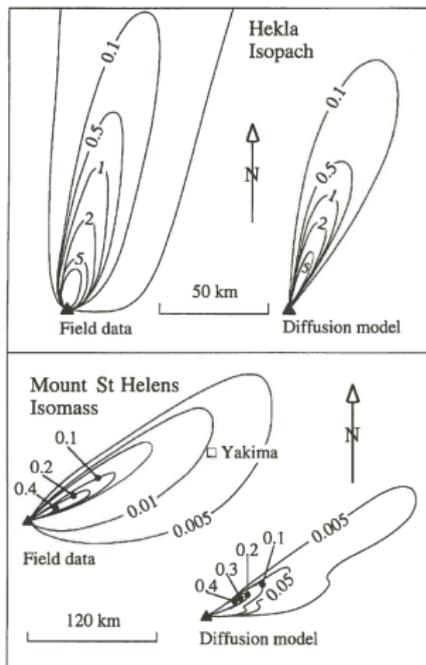
# Particle Settling - Terminal Velocity



Sparks et al., 1997

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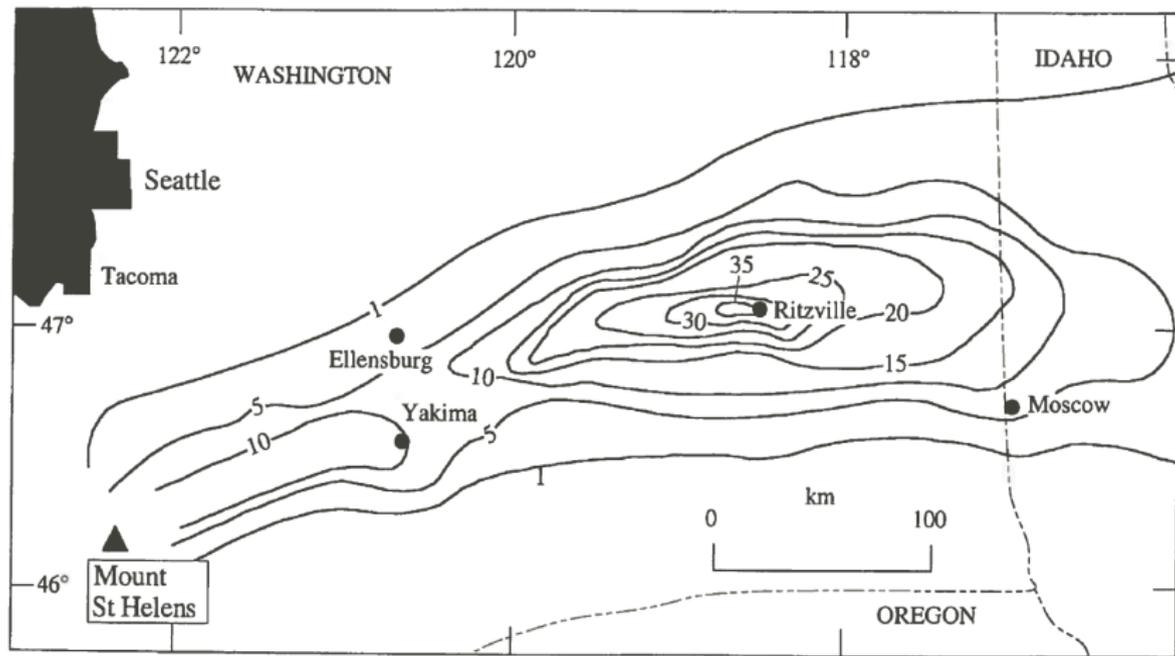
# Particle Settling - Advection / Diffusion modeling



**Figure 14.15** Predicted deposition patterns compared to isopach data for the August 17, 1980 eruption of Hekla, Iceland, and for the July 22, 1980 eruption of Mount St Helens, Washington, USA. The results from the advection/diffusion modelling are in reasonable agreement with field data for these relatively small eruption plumes (from Glaze and Self 1991)

- allow particles to be advected and dispersed in turbulent wind field, 2nd order PDE
- depends on initial concentration of particles in grain size class,
- depends on eddy diffusivity for atmosphere (non-homogeneous)
- depends on changes in source over time

# What's going on & how could this happen?



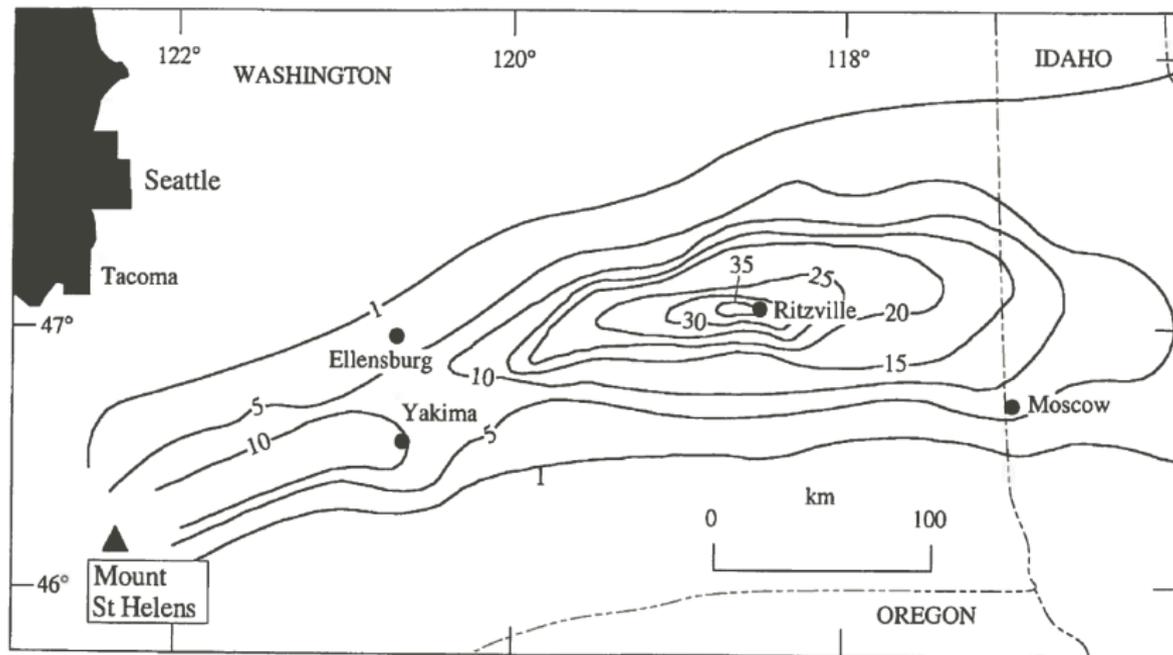
**Figure 13.12** Isopach map of the Plinian fall deposit from the May 18, 1980 eruption of Mount St Helens, modified from Sarna-Wojcicki *et al.* (1981). A second thickness maximum is present in the vicinity of Ritzville, Washington, and has been attributed to premature settling of fine ash by aggregation. Contours are in millimetres

Sparks *et al.*, 1997

# Particle Aggregation

- aggregation of fine ash critical in particle dispersal
- aggregates fall with higher velocity than components (fall out sooner)
- complex grain size distributions
- enhanced thickening of fall deposits
- humidity of plume dictates growth mechanism (dry, accretionary lapilli, mud rain)

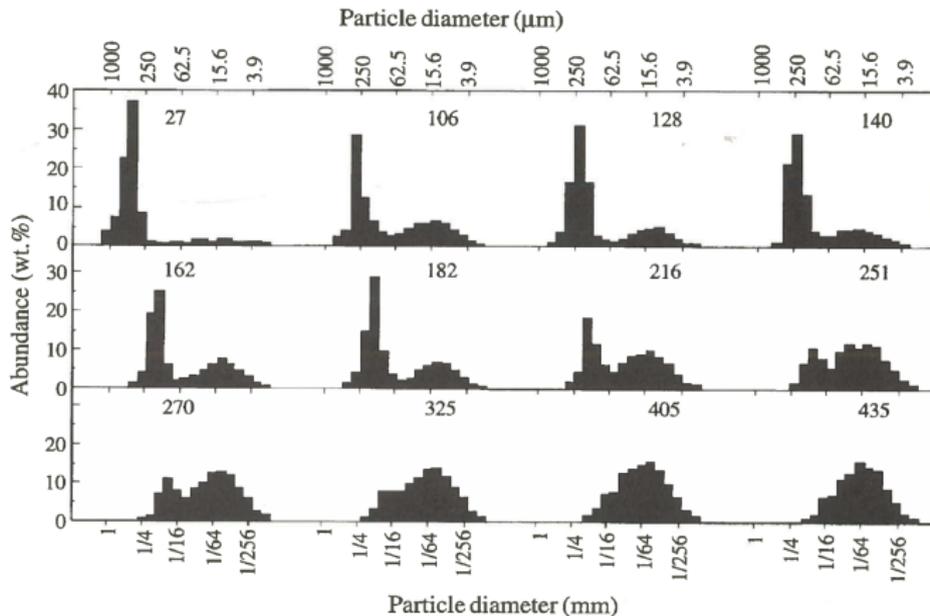
# Anomalous Deposit Thicknesses



**Figure 13.12** Isopach map of the Plinian fall deposit from the May 18, 1980 eruption of Mount St Helens, modified from Sarna-Wojcicki *et al.* (1981). A second thickness maximum is present in the vicinity of Ritzville, Washington, and has been attributed to premature settling of fine ash by aggregation. Contours are in millimetres

Sparks *et al.*, 1997

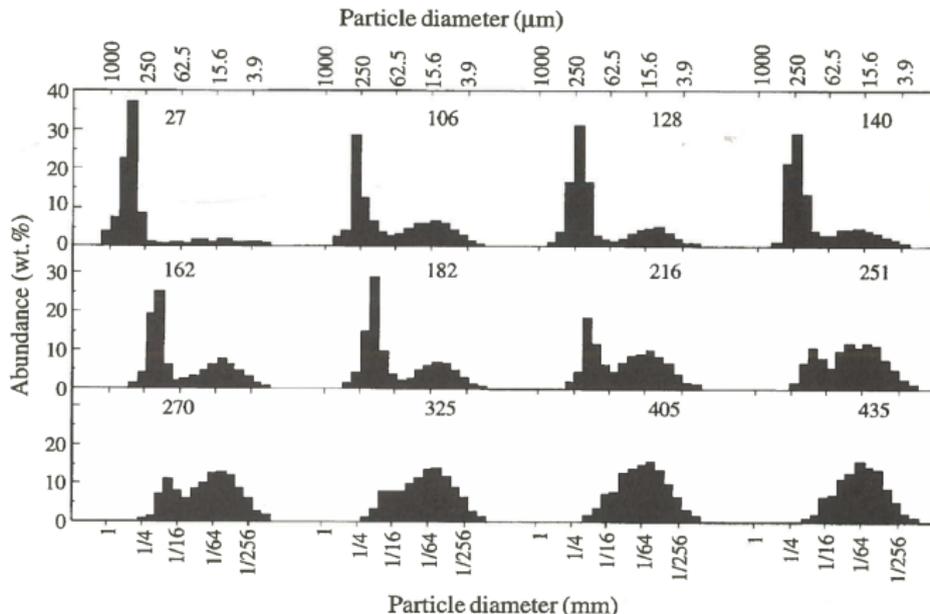
# Particle Size Distributions



**Figure 16.3** Particle size distribution for the bulk deposit of the May 18, 1980 Mount St Helens fall deposit along the dispersal axis. Numbers above histograms are distances, in kilometres, along the dispersal axis. (After Carey and Sigurdsson 1982, reproduced by permission of the American Geophysical Union.)

*Sparks et al., 1997*

# Particle Size Distributions

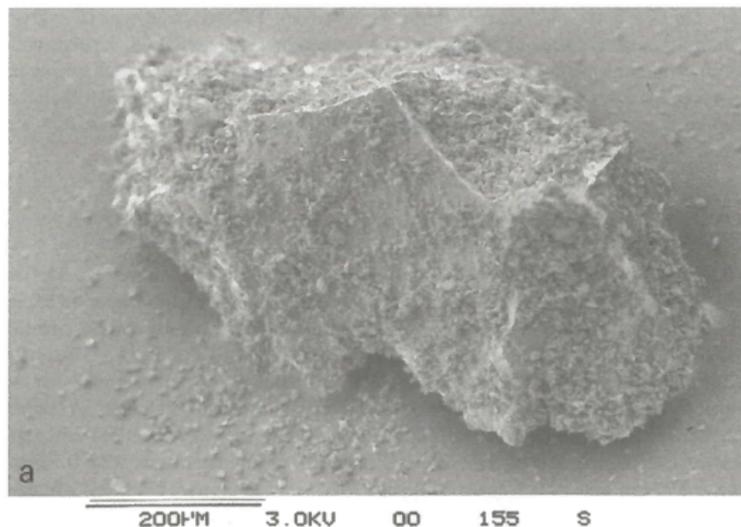


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*Sparks et al., 1997*

polymodal grain size distribution due to deposition of fine material controlled by aggregation

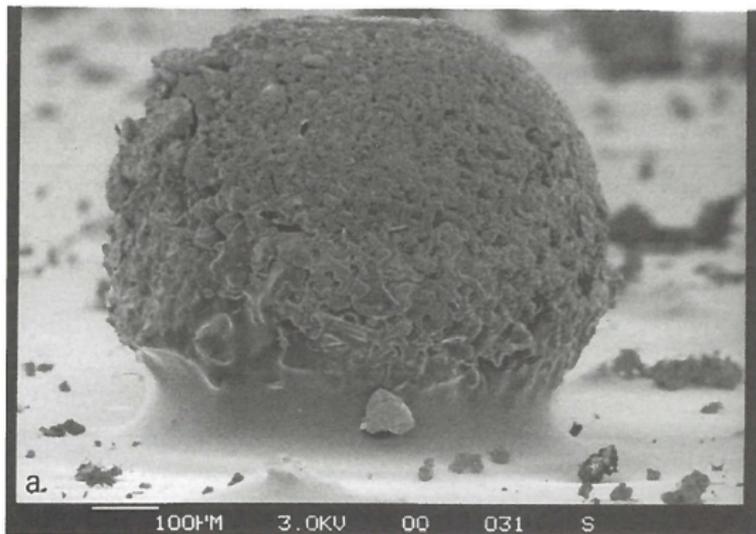
# Dry Aggregates



*Sparks et al., 1997*

Observed at Sakurajima when ground humidity  $< 80\%$

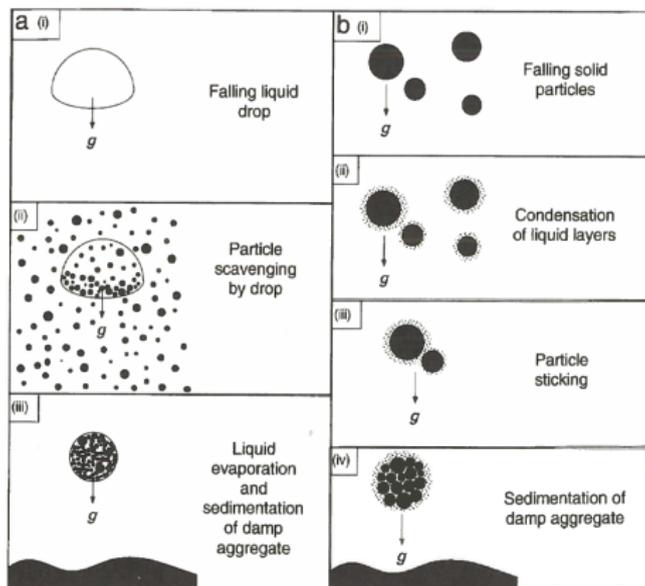
# Accretionary Lapilli



*Sparks et al., 1997*

Observed at Sakurajima when ground humidity > 80%

# Accretionary Lapilli

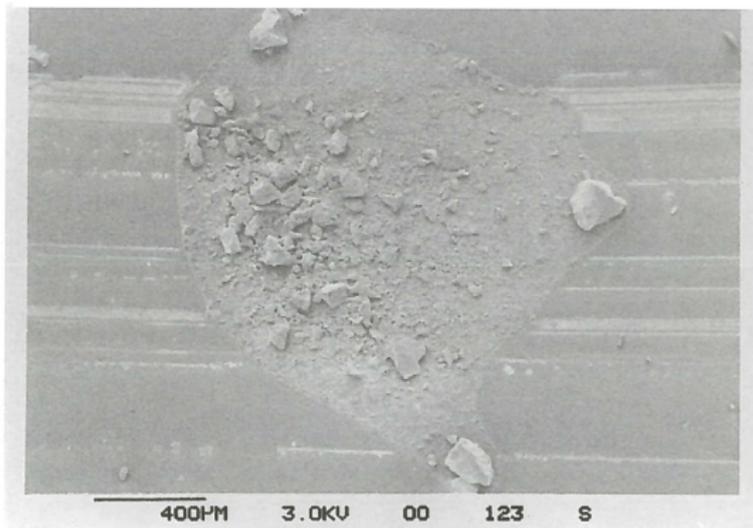


**Figure 16.12** Hypothesis for the evolution of an accretionary lapillus. (a) A liquid drop falling through a plume scavenges particles and partial evaporation of the drop produces a damp aggregate. (b) Particles in a plume, coated with thin liquid layers, collide and coalesce. (After Gilbert and Lane 1994a, reproduced by permission of *Bulletin of Volcanology*.)

Sparks et al., 1997

Observed at Sakurajima when ground humidity  $> 80\%$

# Mud Rain



*Sparks et al., 1997*

# Particle Aggregation

**Table 16.1** Dominant aggregate collision and binding mechanisms, and physical properties of aggregates

	Dry aggregate	Accretionary lapillus	Mud raindrop
<b>Collision mechanisms</b>			
Ambient plume motions	✓	✓	✓
Fall velocity differences	✓	✓	✓
Electrostatic forces	✓		
<b>Binding mechanisms</b>			
Surface tension forces		✓	✓
Secondary mineral/ice crystal growth		✓	
Electrostatic forces	✓		
Van der Waals' forces			
Mechanical interlocking			
<b>Physical properties</b>			
Component particle diameter ( $\mu\text{m}$ )	< 200	< 90	< 200
Aggregate diameter (mm)	< 5	1–50 (but generally 1–10)	< 5
Density ( $\text{kg m}^{-3}$ )	220–1320	1200–1600	1000–1500
Porosity	0.4–0.9	0.3–0.5 (when dry)	0

*Sparks et al., 1997*